

CDSA 2014

**Doppler velocities in kink MHD waves.**

Coupling and Dynamics of the Solar Atmosphere

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FWO Vlaanderen

## 1. Transverse and rotational motions

- **Standing** transverse MHD waves in coronal loops (Schrijver et al. (1999); Aschwanden et al. (1999)); rare events.
- **Propagating** transverse MHD waves (COMP (Tomczyk et al. 2007), SDO/AIA (McIntosh et al. 2011)); are everywhere.
- Transverse waves = **kink** waves ( $m = 1$ ) (Nakariakov et al. (1999)).
- Doppler signals of **torsional motions** and **swaying** motions (De Pontieu et al. (2012))
- Torsional motions = signatures of  $m = 0$  Alfvén waves.
- Axisymmetric ( $m = 0$ ) Alfvén waves and kink ( $m = 1$ ) MHD waves.
- ● ● ● ●
- Kink MHD waves  $\iff$  **transverse + vortical motions**
- Identify vortical motions  $\iff$  Look at Vorticity  $Z$
- Doppler velocities.

## 2. Kink MHD waves on flux tubes

- MHD waves superimposed on an equilibrium magnetic plasma.
- **Equilibrium model** = cylindrical plasma column in static equilibrium.
- Equilibrium magnetic field  $\vec{B}_0 = B_0 \vec{1}_z$ , equilibrium density  $\rho_0(r)$ .
- $\exp(-i\omega t)$ ,  $\exp(i(m\varphi + k_z z))$   $m, k_z =$  **azimuthal and axial wave numbers.**
- $\vec{\xi} =$  displacement,  $P' =$  Eulerian perturbation of total pressure.
- **Vorticity**  $Z = (\nabla \times \vec{\xi}) \cdot \vec{1}_z$

$$Z = (\nabla \times \vec{\xi}) \cdot \vec{1}_z = i \frac{m}{r} P' \frac{d}{dr} \left( \frac{1}{\underbrace{\rho_0(r)}(\omega^2 - \omega_A^2(r))} \right)$$

- **Local Alfvén frequency**  $\omega_A(r)$

$$\omega_A^2(r) = \frac{(\vec{k} \cdot \vec{B})^2}{\mu \rho_0(r)} = k_z^2 v_A^2(r), \quad v_A^2(r) = \frac{B^2}{\mu \rho_0(r)}$$

- $Z$  = measure for local rate of rotation.
- $Z \neq 0$  when?
- Uniform plasma of infinite extent:  $Z \neq 0 \iff$  (bulk) Alfvén waves.
- **Non-uniform plasma:**  $\rho_0(r)$  and  $\omega_A(r)$  depend on position.
- Look back at equation for  $Z$ .
- **Non-uniform plasma:**  $Z \neq 0$
- **Non-uniform plasma: Kink MHD waves contain rotational motions!**
- Details of rotational motions? Where are the transverse motions?
- $\vec{\xi} = ?$

‘Do you realize what this means?’ ‘Yes, I do, but you say it first!’  
*Conversation between Bart and Milhouse*  
 The Simpsons.

### 3. Kink MHD waves: Piece-wise uniform density.

- Piece wise constant  $0 \leq r \leq R : \rho_0 = \rho_i; \quad r \geq R : \rho_0 = \rho_e; \quad \rho_i > \rho_e$ .

$$\rho_e < \rho_i, \quad B_z = \text{constant} \implies \omega_{Ai}^2 = k_z^2 v_{Ai}^2 < \omega_{Ae}^2 = k_z^2 v_{Ae}^2$$

- Wentzel, 1979; Wilson, 1979, 1980; Edwin and Roberts, 1983.
- **Bessel functions**  $J_m(x)$  ( $x = k_i r$ ) ( $0 \leq r \leq R$ ) and  $K_m(y)$  ( $y = k_e r$ ) ( $R \leq r$ ).
- Continuity of  $P'$  and  $\xi_r$  : the dispersion relation.
- Popular special case: Thin tube (TT) approximation  $k_z R \ll 1 = m$ .

$$\omega^2 = \frac{\rho_i \omega_{Ai}^2 + \rho_e \omega_{Ae}^2}{\rho_i + \rho_e} = \omega_k^2, \quad \omega_{Ai}^2 < \omega_k^2 < \omega_{Ae}^2$$

- Spatial eigenfunctions

$$\begin{aligned} \xi_{r,i}(r)/R &= C, & \xi_{\varphi,i}(r)/R &= i C, & 0 \leq r \leq R; \\ \xi_{r,e}(r)/R &\sim dK_1(y)/dy, & \xi_{\varphi,e}(r)/R &\sim K_1(y) & R \leq r < \infty \end{aligned}$$

- Spatial eigenfunctions, general case: Use  $J_1(x)$  and  $K_1(y)$
- $\xi_\varphi$  is discontinuous at  $r = R$  due to change of sign of  $\omega^2 - \omega_A^2$ .

$$\rho_0(\omega^2 - \omega_A^2) \xi_\varphi = \frac{im}{r} P'$$

- **Vorticity  $Z$**

$$Z \sim \delta(x - 1), \quad x = r/R$$

is concentrated at  $r = R$  due to discontinuity in  $\xi_\varphi$ .

- **Rotational motions!**
- Piece wise uniform density: instructive but pathological.

The world is full of obvious things which nobody by any chance ever observes.  
*Sherlock Holmes*, The Hound of the Baskervilles.  
 Sir Arthur Cannon Doyle

#### 4. Kink MHD waves: nonuniform density.

- Continuous variation of density from  $\rho_i$  to  $\rho_e$  in  $[R - l/2, R + l/2]$ .
- Density  $\rho(r)$  decreases from  $\rho_i$  to  $\rho_e$ .
- Alfvén frequency  $\omega_A(r)$  increases from  $\omega_{Ai}$  to  $\omega_{Ae}$ .
- Recall

$$\omega_{Ai}^2 < \omega_{kw}^2 \cong \frac{\rho_i \omega_{Ai}^2 + \rho_e \omega_{Ae}^2}{\rho_i + \rho_e} < \omega_{Ae}^2$$

- Point of resonance  $r_A \in ]R - l/2, R + l/2[$  where

$$\omega_{kw}^2 = \omega_A^2(r_A)$$

- What about vorticity  $Z$ ?

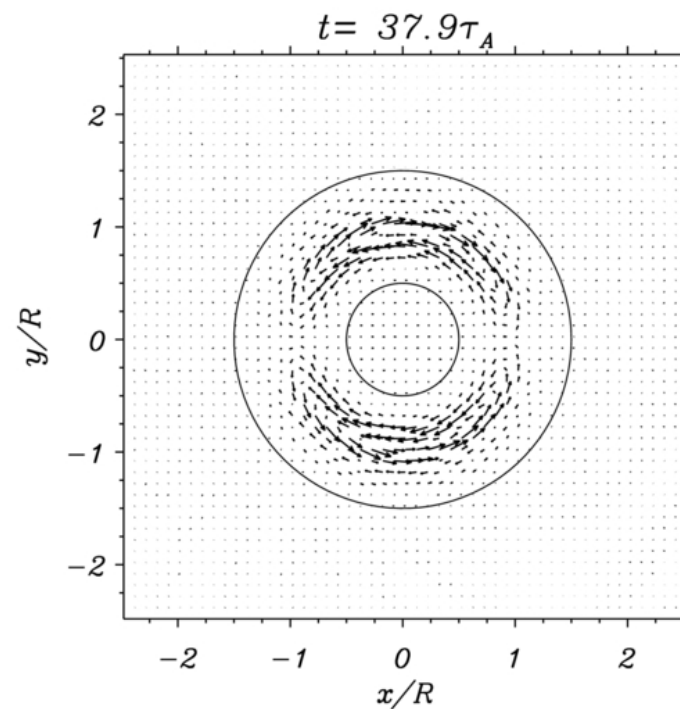
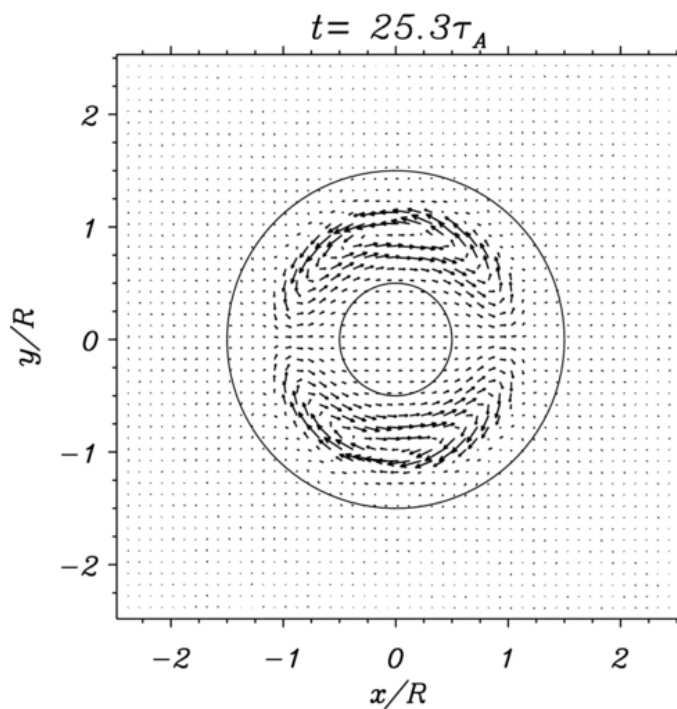
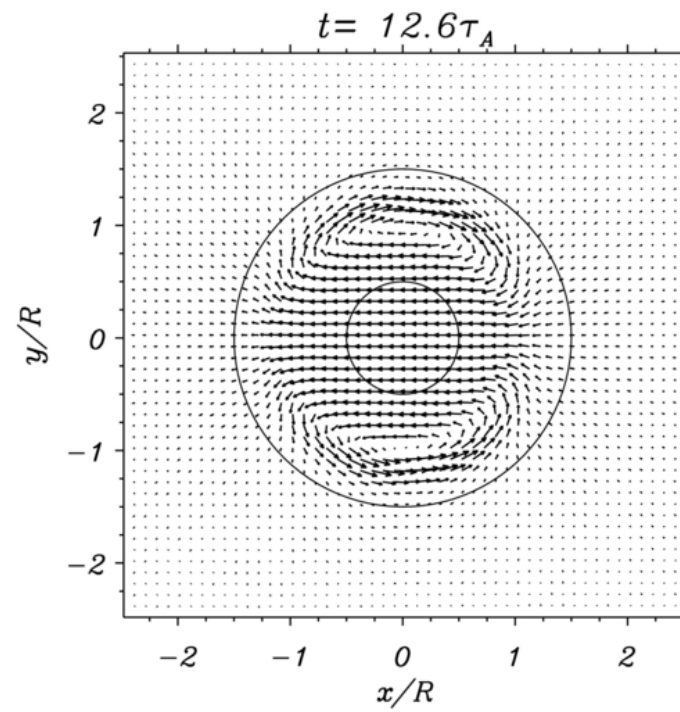
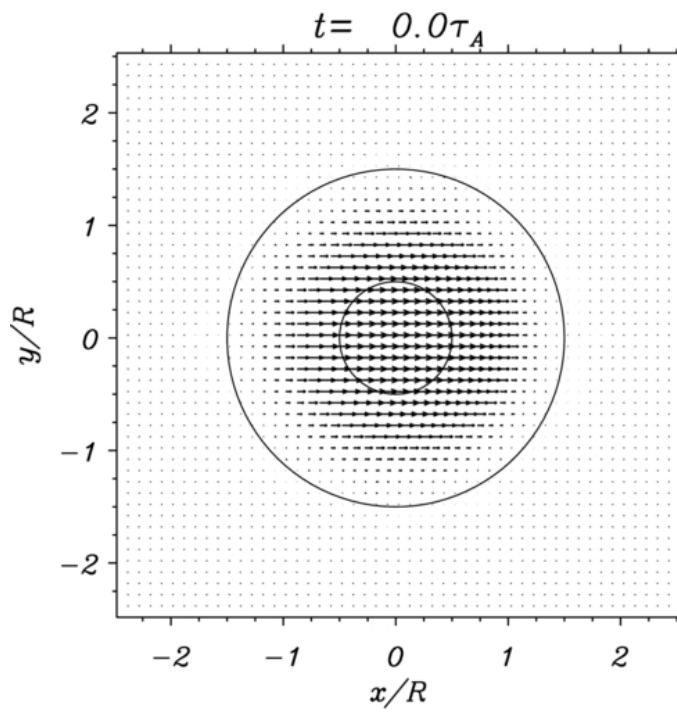
$$Z = (\nabla \times \vec{\xi}) \cdot \vec{1}_z = i \frac{m}{r} P' \frac{d}{dr} \left( \frac{1}{\rho_0(r)(\omega^2 - \omega_A^2(r))} \right)$$

- $Z \neq 0$  in non-uniform layer.
- Ideal MHD and real frequencies: singular solutions.
- $Z \rightarrow \infty$  in ideal MHD for real frequencies at  $r = r_A$ .
- Damped MHD quasi-modes with complex frequencies: finite solutions (Tirry and Goossens, 1996; Van Doorselaere et al. 2004; Soler et al. 2013).
- $Z$  large but finite for damped kink waves.
- At  $r = r_A$  kink wave  $\approx$  pure local Alfvén wave.
- Strong rotational motions in non-uniform plasma in vicinity of  $r = r_A$ .
- Kink wave is Alfvénic in non-uniform plasma.
- Time evolution of kink wave in non-uniform plasma (Terradas et al., 2008; talk J. Terradas at BUKS 2009, Leuven).

”Singularity is invariably a clue.”

*Sherlock Holmes*, The Boscombe Valley Mystery.

A. Cannon Doyle.



## 5. Transverse and rotational motions.

- **First attempt: piece wise uniform density and thin tube approximation.**
- Introduce azimuthal dependency.

- $$0 \leq r \leq R : \frac{\vec{\xi}_h(r, \varphi)}{R} = C (\cos \varphi \vec{1}_r - \sin \varphi \vec{1}_\varphi)$$

- Dependency on  $z$  :  $\exp(ik_z z)$
- Jump in plane  $z = z_0$  and define a system of Cartesian coordinates

$$x = r \cos \varphi, \quad y = r \sin \varphi$$
$$\vec{1}_r = \vec{1}_x \cos \varphi + \vec{1}_y \sin \varphi, \quad \vec{1}_\varphi = -\vec{1}_x \sin \varphi + \vec{1}_y \cos \varphi$$

- From polar to Cartesian coordinates

$$\vec{\xi} = \begin{bmatrix} \xi_x \\ \xi_y \end{bmatrix} = C \begin{bmatrix} 1 \\ 0 \end{bmatrix} = \vec{\xi}_{TR}$$

- **Uniform motion of the entire internal plasma along the  $x$ -axis.**

- How did we get the uniform motion along the  $x$ -axis
- Necessary conditions: uniform, thin tube, piece-wise constant density.
- $0 \leq r \leq R$ . What happens outside tube  $r \geq R$ ?

• **Remember strong rotational motions in non-uniform tube .**

- $\xi_r$  and  $\xi_\varphi$  have unequal amplitudes .

•

$$0 \leq r \leq R: \frac{\vec{\xi}_h(r, \varphi)}{R} = C_{TR} \cos \varphi \vec{1}_r - C_{AZ} \sin \varphi \vec{1}_\varphi, \quad C_{TR} \neq C_{AZ}$$

- Again from polar to Cartesian coordinates

$$\vec{\xi} = \begin{bmatrix} \xi_x \\ \xi_y \end{bmatrix} = C_{TR} \begin{bmatrix} 1 \\ 0 \end{bmatrix} + C_{ROT} \sin \varphi \begin{bmatrix} \sin \varphi \\ -\cos \varphi \end{bmatrix}$$

$$C_{ROT} = C_{AZ} - C_{TR}$$

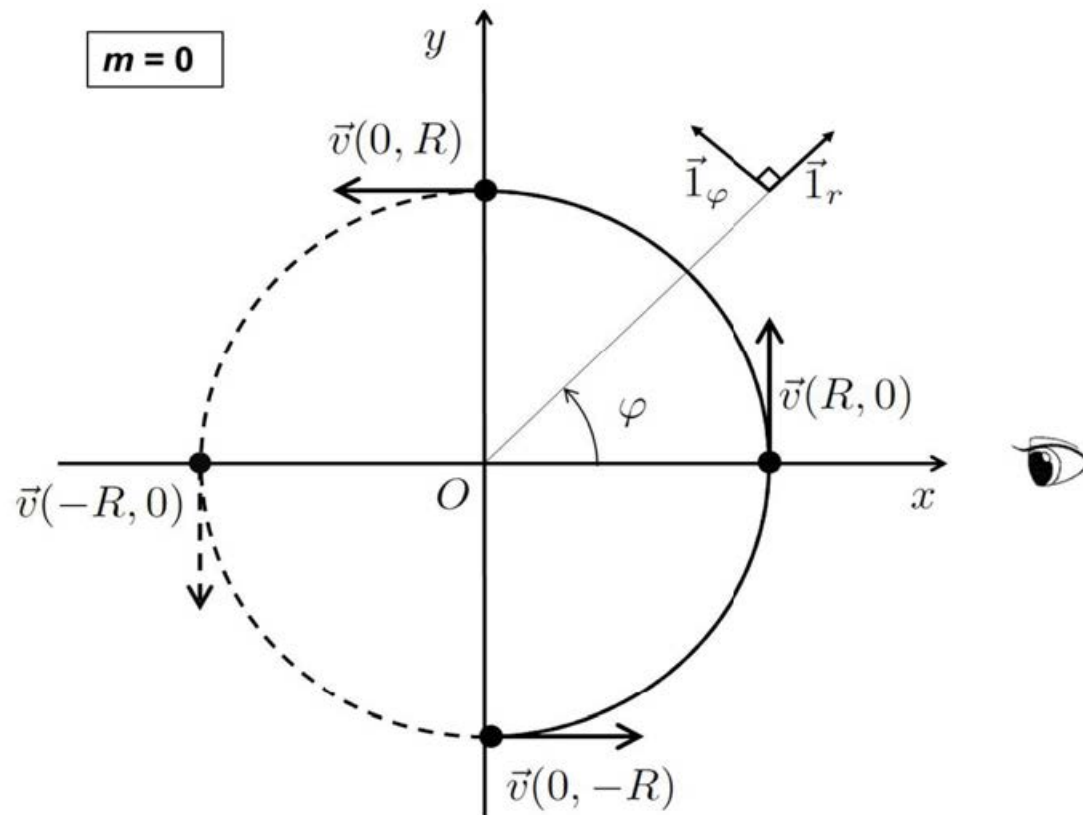
- $\xi_x = \text{constant part} + \text{variable part } \varphi$ .

## 6. Doppler velocities due to axi-symmetric Alfvén waves.

- Velocity in axi-symmetric ( $m = 0$ ) Alfvén wave

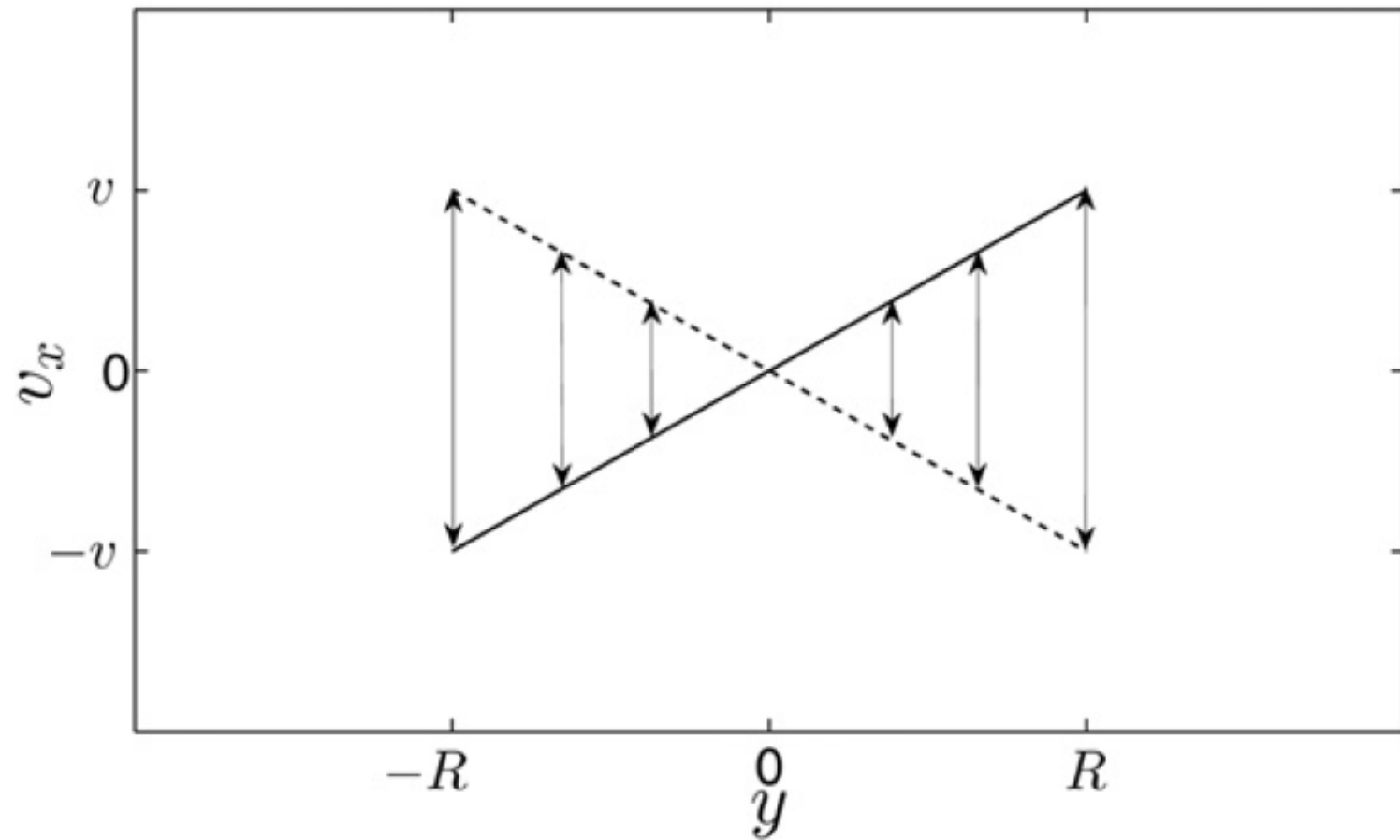
$$\vec{v} = v \vec{I}_\varphi$$

- Take  $x$ -axis in the direction of observer.
- $y$ -axis is across the flux tube



- Component of velocity in the direction of the observer, i.e. along  $x$ -axis

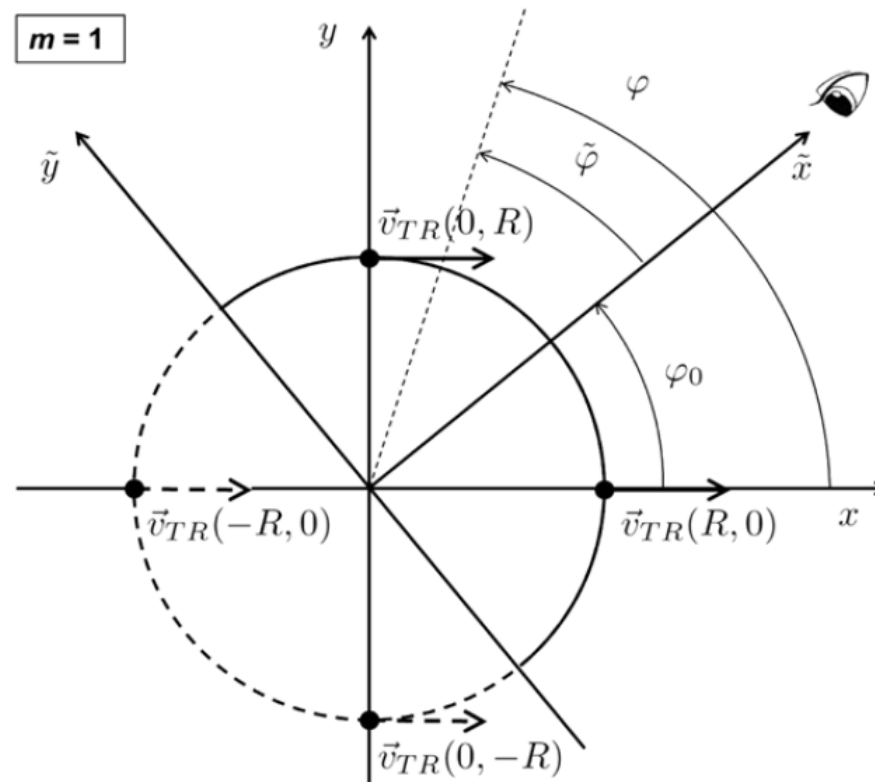
$$v_x = -v \frac{y}{R}$$



- Linear variation.

## 7. Doppler displacements due to kink motions across loop.

- Remember translational motion along  $x$ -axis,  $\varphi = 0$ .
- Line of sight: perpendicular to  $z$ -axis and makes angle  $\varphi_0$  with  $x$ -axis.
- Rotate old system of Cartesian coordinates  $(x, y)$  over angle  $\varphi_0$
- New system of Cartesian coordinates  $(\tilde{x}, \tilde{y})$



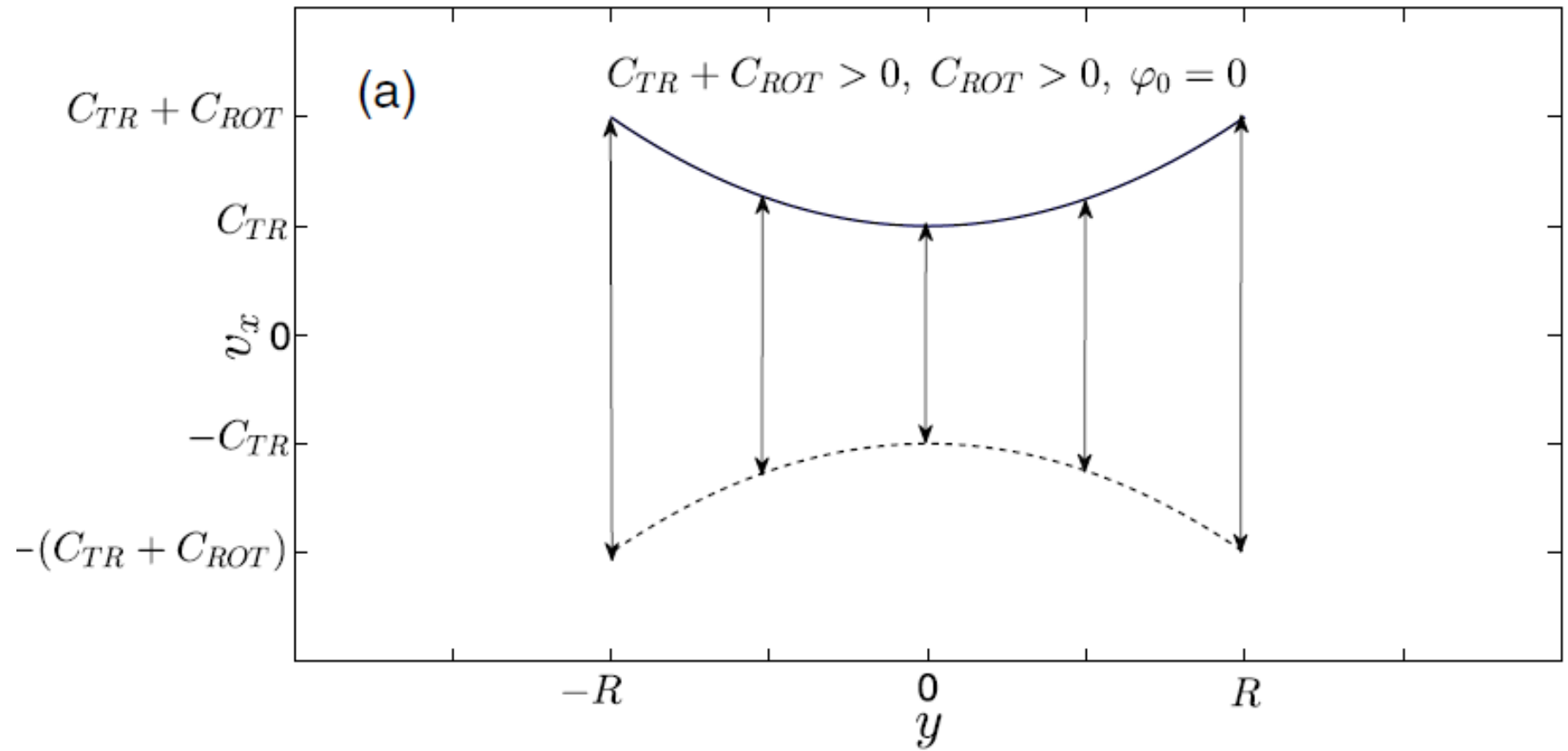
- $\tilde{x}$ -axis is towards the observer,  $\tilde{y}$ -axis is across the flux tube.
- Measure velocity in unit  $\omega R$
- **Purely transverse motion:**  $C_{AZ} = C_{TR}$ ,  $C_{ROT} = 0$
- Component of velocity in the direction of the observer, i.e. along  $\tilde{x}$ -axis

$$v_{\tilde{x}} = C_{TR} \cos \varphi_0$$

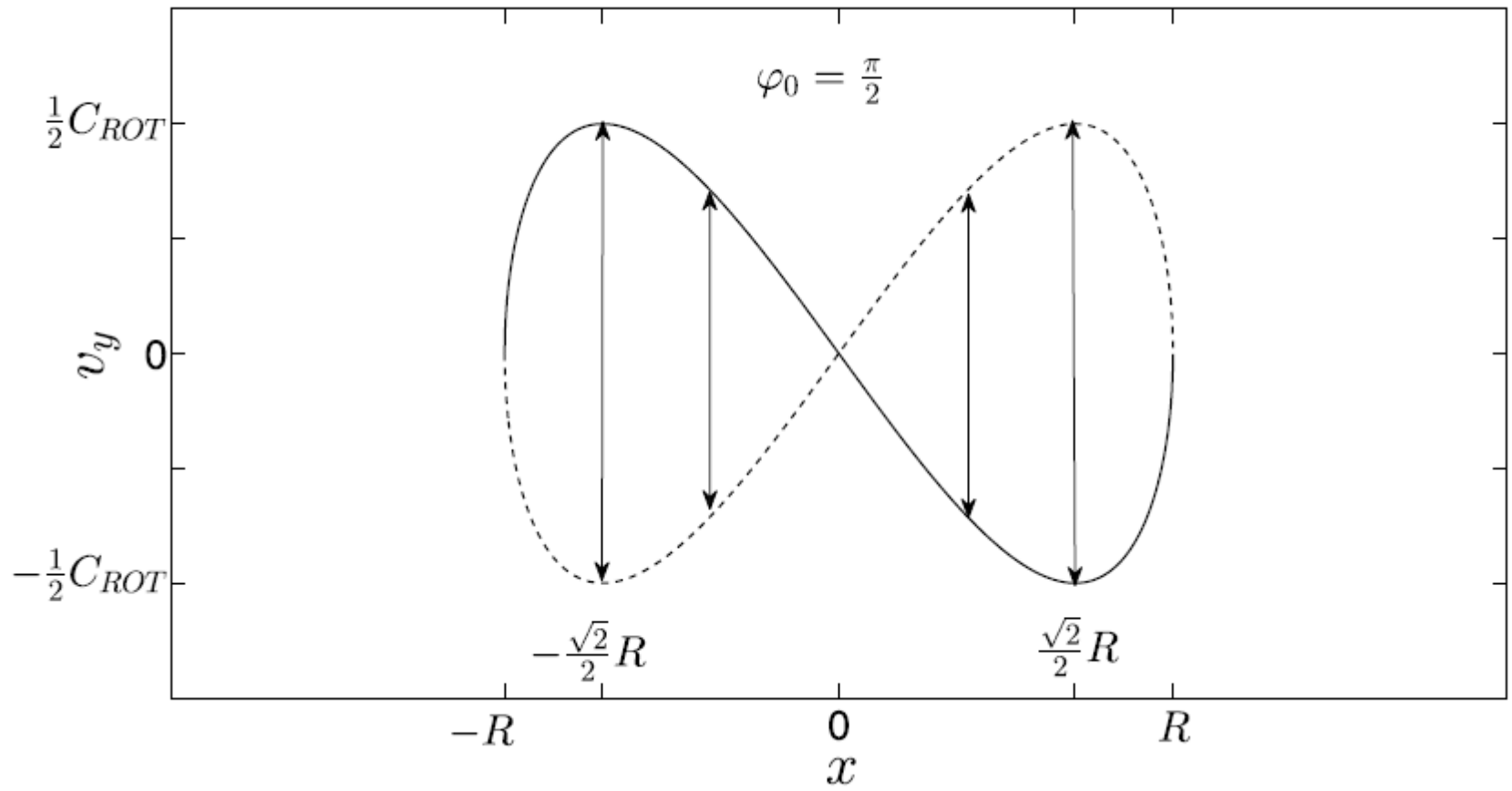
- Constant Doppler velocity across flux tube.
- Role of  $\varphi_0$
- Case  $C_{AZ} \neq C_{TR}$ ,  $C_{ROT} \neq 0$
- Component of velocity in the direction of the observer, i.e. along  $\tilde{x}$ -axis

$$v_{\tilde{x}} = \underbrace{\left\{ C_{TR} + C_{ROT} (\tilde{y}/R)^2 \right\}}_I \cos \varphi_0 + \underbrace{C_{ROT} (\tilde{y}/R) \left( 1 - (\tilde{y}/R)^2 \right)^{1/2}}_{II} \sin \varphi_0$$

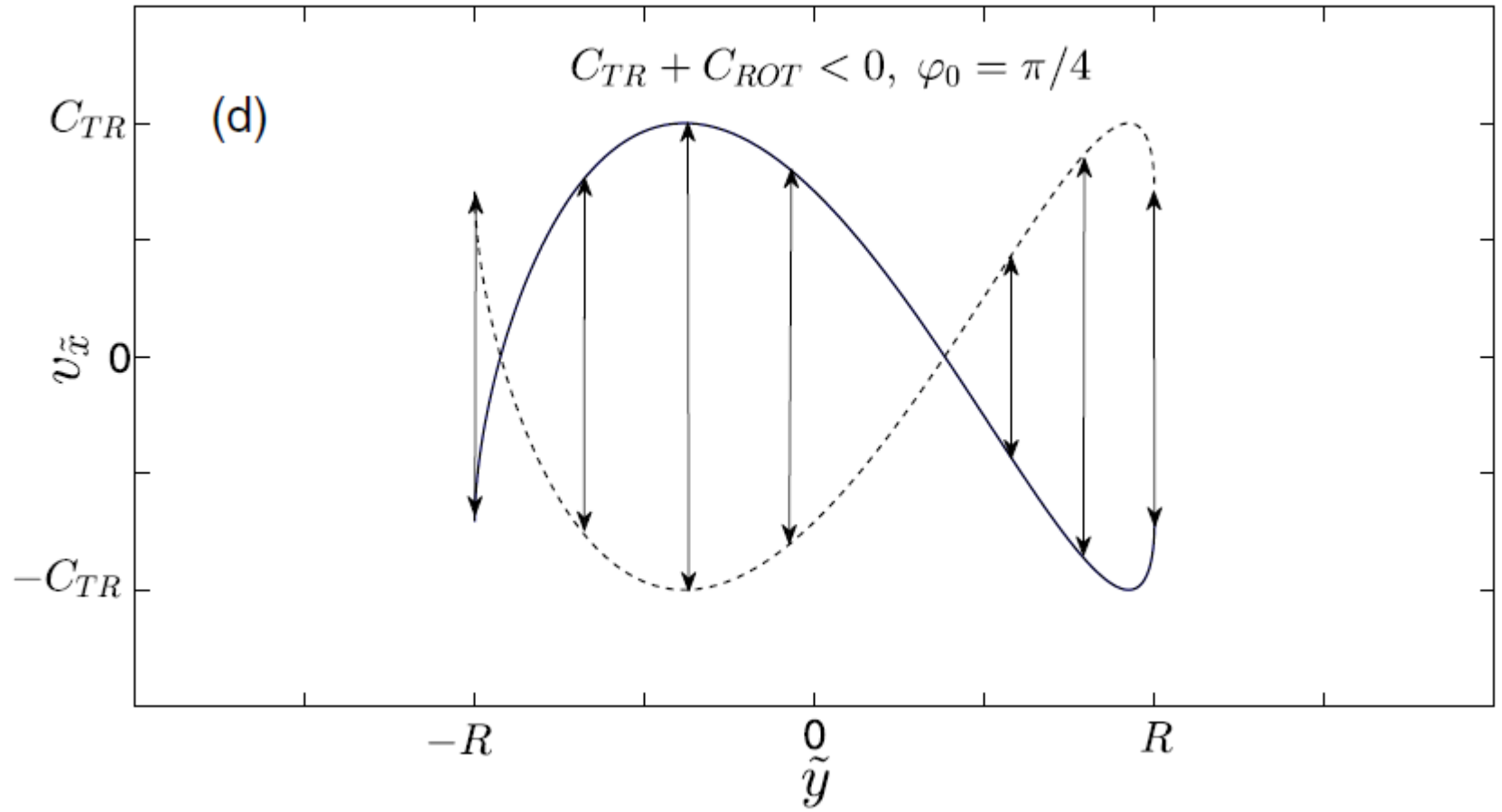
- $I \varphi_0 = 0$ : Parabolic variation (GV : horseshoes.)



- II  $\varphi_0 = \pi/2$

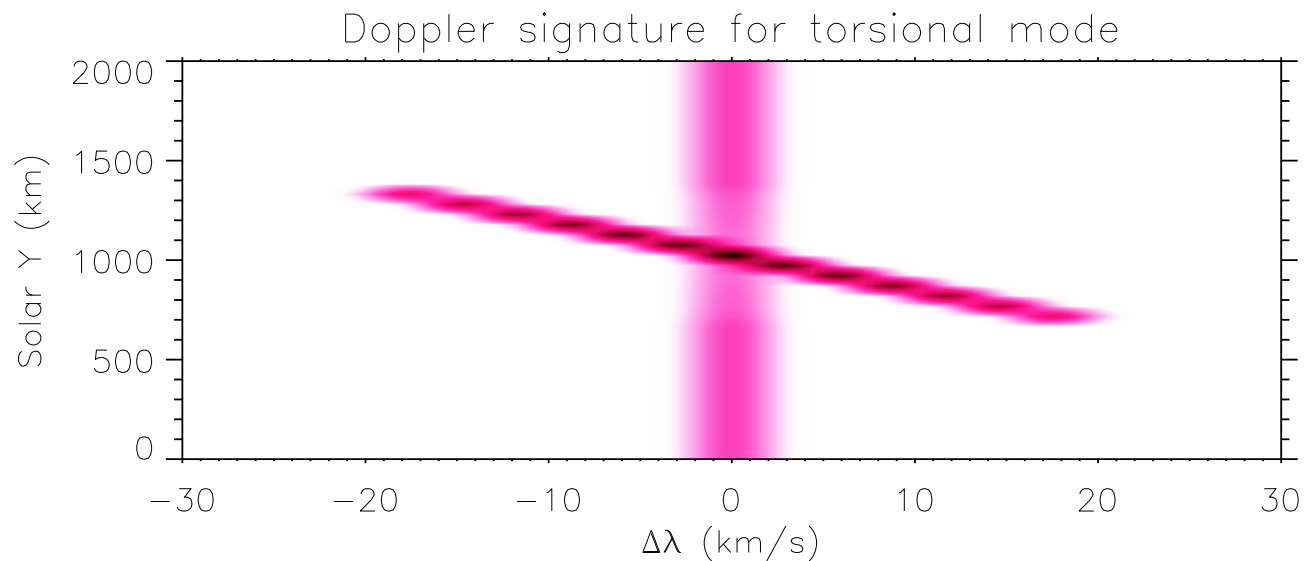


• III General  $\varphi_0$



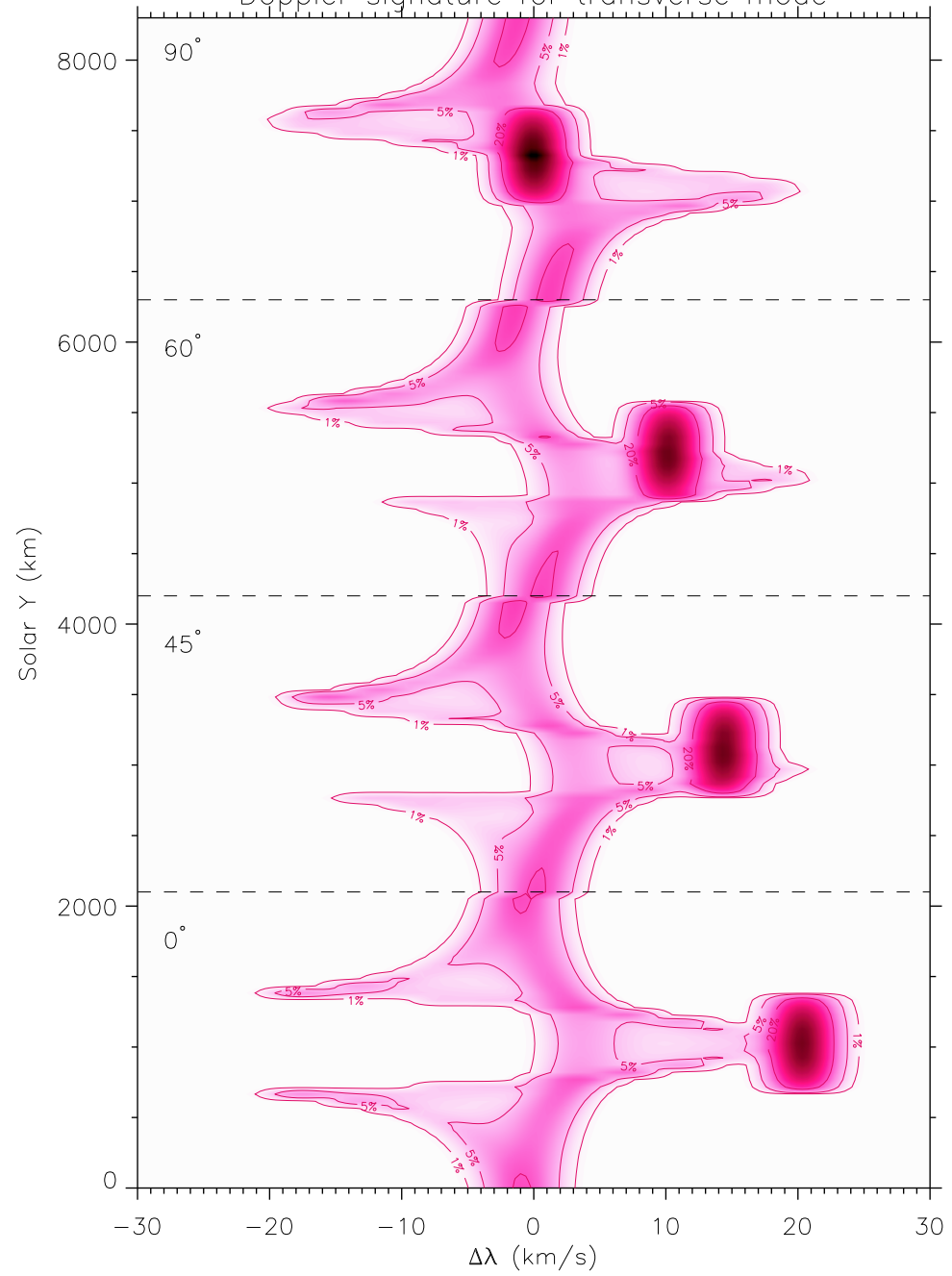
## 8. Integrated intensities for axi-symmetric Alfvén waves and kink waves.

- Details: consult TVD.
- Integrated intensity for the axi-symmetric Alfvén wave.

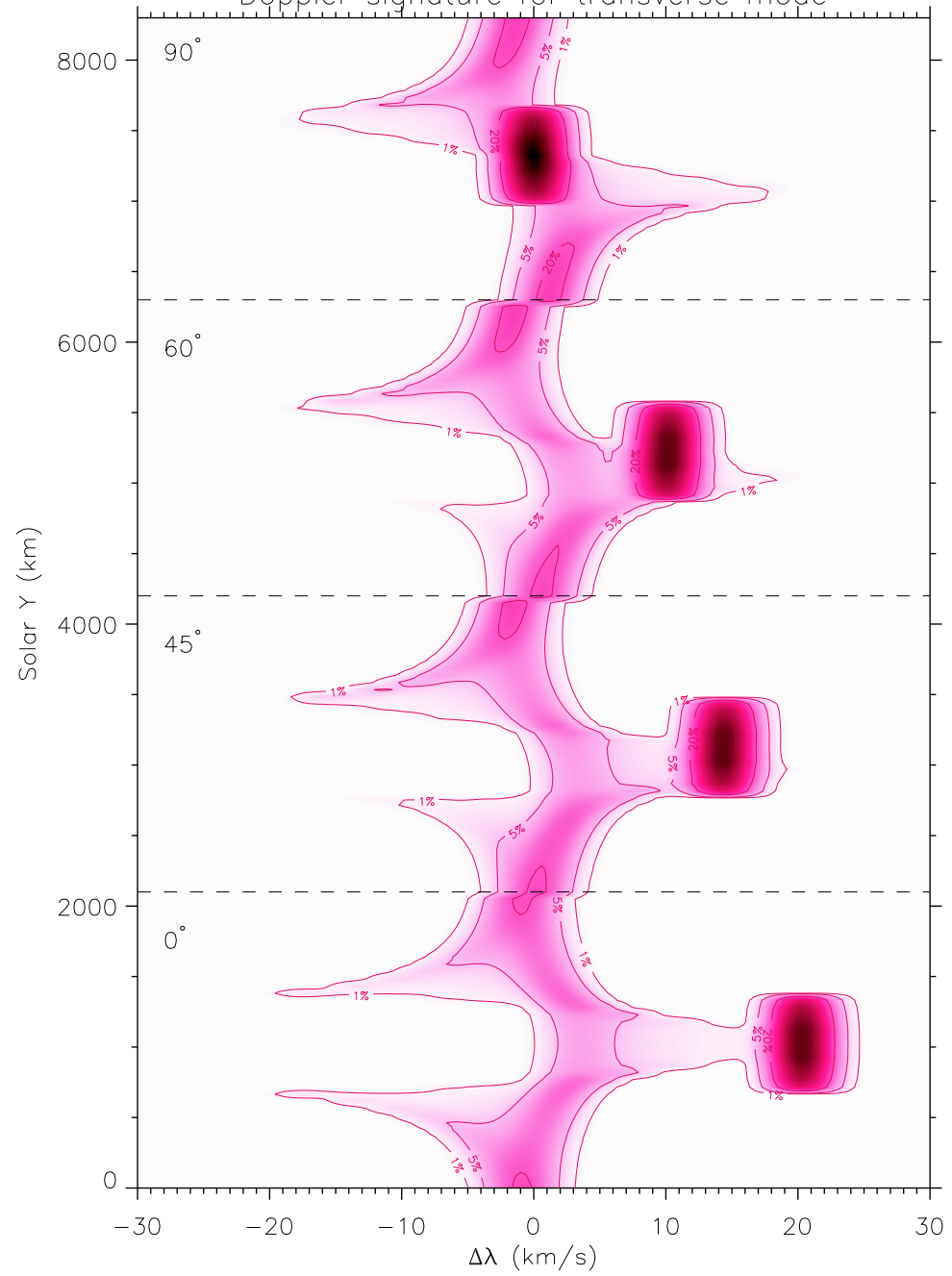


- Result for AW is independent of angle  $\varphi_0$
- Next 2 figures: Integrated intensity for the **kink waves**.
- Piece wise constant density  $\rho_i/\rho_e = 10$
- Smooth density profile  $l/R = 1$

Doppler signature for transverse mode



Doppler signature for transverse mode



## 9. Conclusions

“Has anything escaped me?” I asked with some self-importance.

“I trust that there is nothing of consequence which I have overlooked?”

“I am afraid my dear Watson, that most of your conclusions were erroneous.”

*The Hound of the Baskervilles.*

A. Conan Doyle.

- Kink waves in a non-uniform plasma:  $Z \neq 0$ .
- Motion of plasma = translational motion + rotation.
- Variation of  $v_{\tilde{x}}$  across loop = linear for axi-symmetric Alfvén wave.
- $v_{\tilde{x}}$  = is constant for special case of uniform translation in kink wave.
- Variation of  $v_{\tilde{x}}$  across loop = horseshoe + compressed sinus.
- Integrated intensities for axi-symmetric Alfvén wave and kink waves.

## References

- Goossens, M., Soler, R., Terradas, J., M., Van Doorselaere, T., & Verth, G. 2014, ApJ, 788:9